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SUPERNOVA 2006ao

T. Boles, Coddensham, England, reports the discovery of an apparent supernova on unfiltered CCD images taken on Mar. 1.069 (at mag 17.6) and 1.810 UT (at mag 17.4) with a 0.35-m reflector. The new object is located at $\alpha = 15^{\text{h}}59^{\text{m}}19^{\text{s}}25$, $\delta = +65^{\circ}13'43''.6$ (equinox 2000.0; average position from ten images on two nights), which is $\approx 3''.1$ east and $13''.3$ south of the center of the nearest galaxy, which itself is within a small cluster. Nothing is present at this location on Boles' images from 2005 July 12 and May 11 (limiting mag 19.5) or on Digitized Sky Survey plates from 1994 May 31 (limiting red mag 20.5) and 1992 June 4 (limiting blue mag 21.0).

COMET 73P/SCHWASSMANN-WACHMANN

D. Schleicher, Lowell Observatory, obtained narrowband photometry of comet 73P on Feb. 25 at $r = 1.65$ AU using the Hall 1.1-m telescope at Lowell Observatory, with the following averaged results for component C: $\log Q(\text{OH}) = 27.30$; equivalent $\log Q(\text{water; vectorial}) = 27.33$; $\log Q(\text{CN}) = 24.89$; $\log Q(\text{C}_2) = 24.01$; $\log Af\rho = 1.7$ (cf. *IAUC* 7342). For component B: $\log Q(\text{OH}) = 26.8$; $\log Q(\text{water}) = 26.8$; $\log Af\rho = 1.1$. The relative C₂-to-CN ratio for component C is strongly depleted (by 10 times) with respect to "normal", as defined by A'Hearn *et al.* (1995, *Icarus* **118**, 223) — consistent with the upper-limit result for 73P determined by Fink and Hicks (1996, *Ap.J.* **459**, 729). To within relatively large uncertainties, component B exhibits the same depletion of C₂ as does component C.

V1647 ORIONIS AND ASSOCIATED NEBULA

Further to *IAUC* 8600 and *IBVS* 5661, several observers recently have reported the extreme fading of the nebula associated with V1647 Ori, including B. Reipurth (University of Hawaii), G. M. Ratto (Pisa, Italy), and P. Corelli (Pagnacco, Italy). Corelli writes that his unfiltered CCD frames from Feb. 6 show no nebula to a limiting stellar mag of 18. Reipurth remarks that this appears to confirm its having been a short 'EXor' eruption (cf. *IAUC* 8289).

RS OPHIUCHI

Additional photometry by G. Sostero and E. Guido (cf. *IAUC* 8673): Feb. 21.48 UT, $V = 7.36$, $B-V = +0.70$; 22.48, 7.43, +0.66; 24.48, 7.70, +0.79; 25.48, 7.75, +0.65; 26.47, 7.76, +0.65; 27.48, 8.04, +0.62.