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## COMETS C/2006 A8 AND C/2006 Y10-Y17 (SOHO)

Nine additional near-sun comets have been found on 2006 SOHO website images (cf. IAUC 8844), all being Kreutz sungrazers except apparently for C/2006 Y12. C/2006 A8 was stellar in appearance and very faint (mag  $\approx$  7–7.5) with no tail. All of the remaining objects that were visible in C3 images appeared therein as small and stellar; all of the remaining objects that were visible in C2 images appeared therein as tailless and slightly diffuse, except for C/2006 Y12, which was faint yet more stellar in appearance and unusually small. In the table below, new finder 'AW' = A. Watson.

Comet	2006  UT	$\alpha_{2000}$	$\delta_{2000}$	Inst.	F	MPEC
C/2006 A8	Jan. 10.013	$19^{^{ m h}}\!36\overset{^{ m m}}{.}\!4$	$-24^{\circ}04^{'}$	C3	BZ	2007-L16
C/2006  Y10	Dec. 23.071	$18\ 11.9$	$-26\ 33$	C3/2	TH	2007 - K45
C/2006  Y11	24.821	$18 \ 18.4$	$-25 \ 34$	C3/2	MA	2007 - K45
C/2006  Y12	27.814	$18 \ 31.4$	$-24\ 40$	C2	RK	2007 - K45
C/2006  Y13	28.154	$18 \ 35.4$	$-25\ 53$	C3/2	WX	2007 - K45
C/2006  Y14	28.504	$18 \ 34.6$	$-24\ 38$	C2	RK	2007 - K45
C/2006  Y15	29.179	$18 \ 39.4$	$-25\ 51$	C3/2	WX	2007 - K46
C/2006  Y16	29.413	$18 \ 38.6$	$-24\ 44$	C2	AW	2007 - K46
C/2006  Y17	29.529	$18\ 43.9$	$-25\ 44$	C3	$_{\mathrm{HS}}$	2007 - K46

## V1280 SCORPII

R. C. Puetter, Center for Astrophysics and Space Science, University of California at San Diego; R. J. Rudy, D. K. Lynch, R. W. Russell, S. Mazuk, and R. L. Pearson, The Aerospace Corporation; C. E. Woodward, University of Minnesota; and R. B. Perry, Langley Research Center, NASA, report 0.4- to 2.5- $\mu$ m spectroscopy of this nova (cf. *IAUC* 8803, 8807, 8809, 8812) using the Lick 3-m telescope (+ VNIRIS) on May 6 UT, as well as 0.8-to 5.5- $\mu$ m spectroscopy using the Infrared Telescope Facility 3-m telescope (+ SpeX) on May 31. This nova is still in a very low excitation state, showing strong C I lines and no discernable He I emission. The spectrum shows a significantly narrower [N I] line at 1.04  $\mu$ m than the permitted emission lines. The reddening derived from the O I lines, which is due in part to the dust shell, is E(B-V)=1.7.