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INTERNATIONAL ASTRONOMICAL UNION**

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F. Marchis, Carl Sagan Center at the SETI Institute and University of California at Berkeley; and his collaborators, P. Descamps and J. Berthier, Institut de Mécanique Céleste et de Calcul des Ephémérides, Observatoire de Paris; and J. P. Emery, University of Tennessee at Knoxville, report on observations of (216) Kleopatra, a large M-type main-belt asteroid, recorded with the 10-m Keck II telescope and its adaptive-optics (AO) system between Sept. 19.27 and 19.50 UT, when the target was at $r = 1.237$ AU with a phase angle of $7^{\circ}.5$. The direct images recorded with the NIRC2 camera reach an angular resolution up to $0''.032$ (corresponding to a spatial resolution of 29 km) with the J_{cont} filter (central wavelength at $1.21 \mu\text{m}$), permitting confirmation that the primary has the “dog-bone” shape revealed by Ostro *et al.* (2000, *Science* **288**, 836). The new AO data show additionally the presence of a ~ 5 -km-diameter companion at $0''.72$ (projected distance of 650 km) in p.a. 322° , detected in every image recorded over the 5.5-hr baseline. Careful analysis of the last set of data indicates the presence of a second satellite, smaller (~ 3 km) and located closer to the primary at an apparent distance of 380 km ($0''.42$) in p.a. 333° . With (87) Sylvia (cf. *IAUC* 8582) and (45) Eugenia (cf. *IAUC* 8817), this is the third multiple system with two small-km-sized satellites to be discovered in the main asteroid belt.

COMET P/2008 R1 (GARRADD)

Improved, elliptical orbital elements for this comet (cf. *IAUC* 8969) from *MPEC* 2008-S46:

$$\left. \begin{array}{ll} T = 2008 \text{ July } 25.200 \text{ TT} & \omega = 256^{\circ}.436 \\ e = 0.34203 & \Omega = 52.018 \\ q = 1.79261 \text{ AU} & i = 15.899 \end{array} \right\} 2000.0$$

$$a = 2.72444 \text{ AU} \quad n^{\circ} = 0.219174 \quad P = 4.50 \text{ years}$$

COMET 205P/GIACOBINI

MPEC 2008-S47 contains astrometry and ephemerides for components A, B, and C (cf. *IAUC* 8978), and also revised orbital elements (with non-gravitational parameters) satisfying the primary component during 1896–2008. Visual total-magnitude estimates for component A: Sept. 11.07 UT, 12.7 (J. J. Gonzalez, Palencia, Spain, 20-cm reflector); 16.96, 12.9 (M. Goiato, Araçatuba, Brazil, 22-cm reflector); 19.83, 12.2 (Gonzalez, León).