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INTERNATIONAL ASTRONOMICAL UNION

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 URL <http://www.cfa.harvard.edu/iau/cbat.html> ISSN 0081-0304
 Phone 617-495-7440/7244/7444 (for emergency use only)

COMET P/2008 T4 (HILL)

R. E. Hill reports his discovery of a comet on CCD images taken with the 0.68-m Schmidt telescope in the course of the Catalina Sky Survey (discovery observation tabulated below), a stack of four 30-s exposures showing a diffuse coma 10'' across and a well-condensed nuclear condensation with a broad tail 20'' long in p.a. $\sim 210^\circ$. Images obtained by A. Boattini with the Mt. Lemmon 1.5-m reflector on Oct. 8.44–8.45 UT show a well-condensed coma with diameter 10'' and a broad tail $\sim 30''$ long in p.a. 245° .

| 2008 UT | α_{2000} | δ_{2000} | Mag. |
|--------------|--|-----------------|------|
| Oct. 8.38711 | 1 ^h 52 ^m 52 ^s .10 | +7°39'25''.5 | 17.9 |

The available astrometry (including predisccovery observations from Sept. 28–30), the following preliminary elliptical orbital elements, and an ephemeris appear on *MPEC* 2008-T90.

$$\left. \begin{array}{l}
 T = 2008 \text{ Sept. } 21.832 \text{ TT} \\
 e = 0.44203 \\
 q = 2.44897 \text{ AU} \\
 a = 4.38908 \text{ AU}
 \end{array} \right\} 2000.0$$

$$\left. \begin{array}{l}
 \omega = 331.404 \\
 \Omega = 46.835 \\
 i = 5.540 \\
 n^\circ = 0.107188 \\
 P = 9.20 \text{ years}
 \end{array} \right\}$$

2008 TC₃

J. Borovicka, Astronomical Institute, Czech Academy of Sciences, reports that Z. Charvat (Czech Hydrometeorological Institute) has noticed a bright spot that likely corresponds to the atmospheric entry of 2008 TC₃ (cf. *IAUC* 8990) over northern Sudan on images taken by the weather satellite Meteosat 8 around Oct. 7^d02^h45^m UT; the spot is apparent in all twelve satellite spectral channels, spanning wavelengths 0.5–14 μm . Since the satellite takes images in scanning mode, it takes ~ 5 min to obtain one image; consequently, the exact time of the spot's appearance cannot be inferred easily from the image. The spot is, however, not present in the images taken at nominal times 2^h40^m and 2^h50^m UT. The geographical coordinates of the spot in the visual and near-infrared channels are $\lambda = 32^\circ 16'$ east, $\beta = +20^\circ 97'$ (assuming that the source of light is at sea level). The HRV channel shows an apparent tail ~ 3 km long toward the west-northwest. The infrared channels (displaced in the instrument focal plane and likely scanned the region 1.8 s later) show the spot at $\lambda = 32^\circ 37'$ east, $\beta = +20^\circ 89'$. The actual coordinates may be slightly southwest of these positions after correcting for the fireball's altitude (~ 30 and 22 km).